



1 Development of a new gas flaring emission data set for southern West Africa

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8 HIGHLIGHTS

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- 10 - Development of a new gas flaring emission parameterization for air pollution modeling.
- 11 - Combination of remote sensing observation and physical based combustion calculation.
- 12 - Application to the significant gas flaring region southern West Africa.
- 13 - Comprehensive assessing of the parameterization uncertainties.
- 14 - Comparison with existing gas flaring emission inventories.

15

16 Keywords:

17

18 Gas flaring

19 Emission parameterization

20 Emission uncertainty

21 Pollution modeling

22 Carbon dioxide

23

24 ABSTRACT

25

26 A new gas flaring emission parameterization has been developed which combines remote sensing
27 observations using VIIRS nighttime data with combustion equations. The parameterization has been
28 applied to southern West Africa, including the Niger Delta as a region which is highly exposed to gas
29 flaring. Two two-month datasets for June-July 2014 and 2015 were created. The parameterization
30 delivers emissions of CO, CO₂, NO and NO₂. A flaring climatology for both time periods has been
31 derived. The uncertainties owing to cloud cover, parameter selection, natural gas composition and
32 the interannual differences are assessed. Largest uncertainties in the emission estimation are linked
33 to the parameter selection. By using remote sensing cloud cover observations, a correction factor for
34 the climatology was established to consider the effect of flares masked by clouds. It can be shown
35 that the flaring emissions in SWA have significantly decreased by 30% from 2014 to 2015. Existing
36 emission inventories were used for validation. CO₂ emissions with the estimated uncertainty in
37 brackets of 8 (¹²/₂) Tg y⁻¹ for 2014 and 5 (⁷/₁) Tg y⁻¹ for 2015 are derived. The flaring emission
38 estimation within this study for June-July 2014 is in the same order of magnitude compared to
39 existing emission inventories. For the same period in 2015 the emission estimation is one order of
40 magnitude smaller in comparison to existing inventories. The deviations might be attributed to
41 uncertainties in the derived flare gas flow rate, the decreasing trend in gas flaring or inconsistent
42 emission sector definitions. The parameterization source code is available as a package of R scripts.

43

44 1. Introduction

45

46 Gas flaring is a globally used method to dispose flammable, toxic or corrosive vapors to less reactive
47 compounds at oil production sites and refineries. In regions of insufficient transportation
48 infrastructure or missing consumers, flaring is also commonly applied.



49 CDIAAC (2015a) estimated the global gas flaring emission of carbon dioxide to 267.7 million tons
50 (0.83% of total emissions) in 2008. Flaring and venting of gas significantly contributes to the
51 greenhouse gas emissions and therefore to the global climate change. The five countries with the
52 highest flaring amount in billion cubic meter (bcm) are Russia (35), Nigeria (15), Iran (10), Iraq (10)
53 and USA (5) (World Bank, 2012).

54 In recent time, especially with the development of remote sensing observation techniques (e.g.
55 Elvidge et al. (1997, 2013)), emissions from gas flaring moved in the focus of atmospheric research
56 involving the efforts in reducing the pollution and the waste of resources. The World Bank led the
57 initiatives “Global Gas Flaring Reduction Partnership” (GGFR) and “Zero Routine Flaring by 2030” to
58 promote the efficient use of flare gas.

59 Instead of relying on national statistics of gas production and consumption for estimating the flaring
60 amount, remote sensing techniques can estimate the flaring amount directly via multispectral data
61 (Elvidge et al., 2013). Elvidge et al. (2009) developed a 15 year dataset of global and national gas
62 flaring efficiency from 1994 to 2008 by using data from the Defense Meteorological Satellite Program
63 (DMSP). Doumbia et al. (2014) combined DMSP with emission factors for flaring, to estimate the
64 flaring emissions for SWA. The satellite product Visible Infrared Imaging Radiometer Suite (VIIRS)
65 Nightfire (Elvidge et al., 2013), which is free available as “VIIRS Nightfire Prerun V2.1 Flares only”
66 (VIIRS, 2015) (VNP hereafter), is now the most widely used product to derive flaring emissions from
67 satellite imagery. By using VNP, Zhang et al. (2015) estimated the methane consumption and the
68 release of CO₂ from gas flaring for the northern U.S. which agree with field data within an uncertainty
69 range of $\pm 50\%$.

70 Also in the second largest flaring country Nigeria, the awareness of gas flaring increases. On
71 gasflaretracker.ng the attention of the government, industry and society is called to the flaring
72 problem by interactive maps of flare infrastructure, amounts and costs. The implications of gas
73 flaring in Nigeria are far-reaching. It influences the environment by noise and deterioration of the air
74 quality (Osuji and Awwiri, 2005). Nwankwo and Ogagarue (2011) have measured higher
75 concentrations of heavy metals in surface water of a gas flared environment in Delta State Nigeria.
76 Adverse ecological and bacterial spectrum modifications by gas flaring are indicated by Nwaugo et
77 al. (2006). Gas flaring also causes acid rain which causes economic burden via rapid corrosion of zinc
78 roofs (Ekpoh and Obia, 2010) and causes retardation in crop growth owing to high temperatures
79 (Dung et al., 2008).

80 The project DACCIWA (Dynamics-aerosol-cloud interactions in West Africa, Knippertz et al. (2015))
81 investigates the influence of anthropogenic and natural emissions on the atmospheric composition
82 over SWA, including the flaring hotspot Nigeria, to examine the meteorological and socio-economic
83 effects. To consider the SWA gas flaring emissions (e.g. in an atmospheric model), this study presents
84 a method to derive emission fluxes by combining the state of the art flaring detection VNP and the
85 combustion equations of Ismail and Umukoro (2014) which does not use emission factors. The new
86 parameterization is robust and easy to apply to new research questions according flexibility in the
87 spatiotemporal resolution.

88 The parameterization is presented in Section 2. Results of the application to SWA, including the
89 spatial distribution of gas flaring, the emission estimation and the uncertainty assessment are
90 investigated in Section 3. Section 4 places the emission estimates in the context of existing
91 inventories. The results are summarized and discussed in Section 5.

92

93 **2. Parameterization of gas flaring emissions**

94



95 The new parameterization for gas flaring presented here, is based on VNP (VIIRS Nightfire Prerun
 96 V2.1 Flares only) and the combustion equations of Ismail and Umukoro (2014) (IU14 hereafter).

97

98 2.1 Remote sensing identification of gas flares

99

100 VIIRS (Visible Infrared Imaging Radiometer Suite) is a scanning radiometer for visible and infrared
 101 light on board the sun-synchronous Suomi National Polar-orbiting Partnership weather satellite
 102 (Suomi-NPP) (NASA, 2016). It can detect combustion sources at night (e.g. bush fires or gas flares) by
 103 spectral band M10. To confirm these sources and to eliminate noise, the Day/Night Band (DNB), M7,
 104 M8 and M12 are used in addition. By fitting these measured spectra to the Planck radiation curve,
 105 background and source temperatures can be deduced. VNP is filtered to include only detections with
 106 temperatures between 1600 K and 2000 K, which is believed to be an adequate estimation for
 107 average gas flares. Up to now no atmospheric correction is done (VIIRS, 2015).

108 The data is freely available as daily data from March 2014 to present. The files include among others
 109 the location of the flares, source temperature T_s , radiant heat H and time of observation.

110 For this study we have decided for a two month period of observation. This allows a compilation of a
 111 flaring climatology in terms of the locations and emissions and a robust estimation of uncertainty
 112 owing to cloud coverage and other parameters that have to be prescribed for IU14. We have
 113 selected the month June and July because the gas flaring emission dataset will be used within the
 114 regional online-coupled chemistry model COSMO-ART (Vogel et al., 2009) during the measurement
 115 campaign of the project DACCIWA, which takes place in June/July 2016. This campaign includes
 116 airborne, ground based and remote sensing observations of meteorological conditions and air
 117 pollution characteristics. COSMO-ART is one of the forecasting models of the DACCIWA campaign
 118 and delivers spatiotemporal aerosol/chemistry distributions. The data for 2014 and 2015 are used to
 119 allow also for an interannual comparison, to assess the uncertainty owing to changes in flare
 120 processes (e.g. built-up or dismantling, increase or decrease in combustion).

121 For this study we use location, source temperature and radiant heat for days with sufficient satellite
 122 coverage over the research domain SWA with a focus on the Niger Delta.

123

124 2.2 Emission estimation method

125

126 The principle emission estimation methodology used in this study follows IU14. The gas flaring
 127 emissions are estimated based on combustion equations for incomplete combustion including six
 128 flaring conditions given in Tab. 1. The equations are introduced in detail in IU14 and are therefore
 129 not presented here. This section concentrates on the application of the method of IU14 to the VNP
 130 data and the research domain in SWA.

131

132 **Tab.1.** Reaction types for incomplete combustion of flared gas, depending on availability of sulfur in the flared gas and the
 133 temperature in the combustion zone which determines the formation of NO and NO₂.

134

Reaction type	Sulfur in flared gas	Source temperature (K)	NO _x formation
1	No	< 1200	no
2	Yes	< 1200	no
3	No	$1200 \leq T_s \leq 1600$	only NO
4	Yes	$1200 \leq T_s \leq 1600$	only NO
5	No	> 1600	NO and NO ₂
6	Yes	> 1600	NO and NO ₂

135



136 As input, IU14 needs the natural gas composition C of the fuel input of the flare, the source
137 temperature T_s (temperature in the combustion zone), and the flare characteristics including
138 combustion efficiency η (1 is complete combustion without Carbon monoxide formation) and
139 availability of combustion air δ (above 1 is excess and below 1 is deficiency). In addition we need the
140 flow rate F , the gauge pressure of the fuel gas in the flare p_g , and the fraction of total reaction
141 energy that is radiated f . The value for f is estimated by averaging a table of literature values for f
142 given in Guigard et al. (2000). The IU14 input is summarized in Tab. 2.

143

144

145

Tab.2. Variables and parameters needed for IU14 or for deriving the fluxes of the air pollutants

Parameter	Description	Reference	Unit
C	Natural gas composition	Sonibare and Akeredolu (2004)	%
T_s	Source temperature	VNP (VIIRS, 2015)	K
η	Combustion efficiency	0.8 (IU14)	-
δ	Availability of combustion air	0.95 (IU14)	-
H	Radiant heat	VNP (VIIRS, 2015)	MW
F	Flow rate	VNP (VIIRS, 2015), TA-Luft (1986)	$\text{m}^3 \text{s}^{-1}$
p_g	Gauge pressure	34.475 (API, 2007)	kPa
f	Fraction of radiated heat	0.27 (Guigard et al., 2000)	-

146

147 The natural gas composition is taken from Sonibare and Akeredolu (2004). They have measured the
148 molar composition of Nigerian natural gas in the Niger Delta area for ten gas flow stations. For this
149 study we have calculated the average over these stations and merged the data according their
150 number of carbon atoms (Tab. 3). H_2S fraction is rather low because it was detected only in two out
151 of the ten flow stations.

152

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Tab.3. Molar composition of natural gas in Niger Delta (Nigeria) based on the measurements of Sonibare and Akeredolu (2004), averaged over ten flow station. The hydrocarbons are merged according to the number of C atoms.

Constituent	Fraction (%)
Methan (CH_4)	78.47
Ethan (C_2H_6)	6.16
Propane (C_3H_8)	5.50
Butan (C_4H_{10})	5.19
Pentane (C_5H_{12})	3.95
Hexane (C_6H_{14})	0.36
Carbon dioxide (CO_2)	0.305
Nitrogen (N_2)	0.06
Hydrogen sulfide (H_2S)	0.005

156

157 The source Temperature T_s is taken from VNP. The combustion efficiency η was set to 0.8 and the
158 availability of combustion air δ to 0.95. IU14 remarked, that the reaction condition for flaring of
159 $\eta \gg 0.5$ and $\delta \geq 0.9$ should be the norm in regions, where the effective utilization of this gas is not
160 available or not economically. Strosher (2000) indicate a combustion efficiency of solution gas at oil-
161 field battery sites between 0.62 and 0.82, and 0.96 for flaring of natural gas in the open atmosphere
162 under turbulent conditions. EPA (1985) shows combustion efficiencies between 0.982 and 1 for
163 measurements on a flare screening facility. Section 3.3.2 will shed light on the uncertainty which
164 arises from η and δ via a parameter sensitivity study. The authors strongly recommend a careful
165 selection of η and δ since unrealistic combinations (e.g. higher combustion efficiencies with rather
166 low availability of combustion air) can lead to negative NO and NO_2 emissions.



167 The flow rate, gauge pressure and fraction of radiated heat are not included in the parameterization
168 of IU14 but are necessary to derive the mass emission rates which can be used as emission data for
169 an atmospheric dispersion model.

170 The flow rate F ($\text{m}^3 \text{s}^{-1}$) is estimated by Eq. 1 (TA-Luft, 1986)

171

$$172 \quad F = M / (1.36 \cdot 10^{-3} (T_S - 283)), \quad (1)$$

173

174 where M is the heat flow in MW and T_S the source temperature in K. We assume that the emitted
175 heat flow M is equal to the total reaction energy of the flare. VNP only detects the energy fraction
176 that is radiated H and not the total energy M . By using the radiant heat H (observed by VNP) and the
177 factor f (fraction of H to the total reaction energy, Guigard et al., 2000), we estimate M as $H \cdot 1/f$.
178 For the source temperature T_S we use the VNP observation.

179 The estimation of the fuel gas density, which is necessary to transform the flow rate F into an
180 emission, is problematic due to the lack of data concerning the technical setup of the SWA flares. We
181 assume that the dominating flare type is a low-pressure single point flare. Bader et al. (2011) pointed
182 out that these flares are the most common flare type for onshore facilities that operate at low
183 pressure (below 10 psi (69 kPa) above ambient pressure) and API (2007) remarks that most subsonic-
184 flare seal drums operate in the range from 0 psi to 5 psi (34 kPa). Therefore we have decided for a
185 gauge pressure p_g of 5 psi (34 kPa) above ambient pressure. Via Eq. 2 we can calculate the fuel gas
186 density ρ_f

187

$$188 \quad \rho_f = p_f / (R / (M_f T_a)), \quad (2)$$

189

190 where p_f is the fuel gas pressure as the sum of ambient pressure (10.1325 kPa, taken as const) and
191 gauge pressure p_g . R is the universal gas constant, M_f the molar mass of the fuel gas and T_a the
192 ambient temperature (293.15 K, taken as const). Finally, the emission E (kg s^{-1}) of a species i is given
193 by

194

$$195 \quad E_i = \frac{m_i}{m_{total}} \rho_f F, \quad (3)$$

196

196 where m_i is the mass of the species i and m_{total} the total mass of the fuel gas, both delivered by the
197 parameterization of IU14.

198 The combustion calculations within IU14 provide the species water, hydrogen, oxygen, nitrogen,
199 carbon dioxide, carbon monoxide, sulfur dioxide, nitrogen oxide and nitrogen dioxide. In the
200 following only CO, SO₂, NO and NO₂ are considered. However, no black carbon or volatile organic
201 compounds (VOCs) are considered by IU14, although they are not negligible. Johnson et al. (2011)
202 estimated the mean black carbon emission for a large-scale flare at a gas plant in Uzbekistan to be
203 7400 g h⁻¹ and Strosher (1996) measured the concentration of predominant VOCs 5 m above the gas
204 flare in Alberta with 458.6 mg m⁻³. However, owing to the missing representation of black carbon and
205 VOCs in IU14, these compounds are not considered in this study.

206 A flaring emission comparison between several days or averaging over a certain period is problematic
207 due to small variances in the VNP locations of the flares. This means even the same flare can be
208 detected on a slightly different position the next day, which makes an emission averaging for every
209 single flare difficult, especially in intensive flare areas. We bypass the problem by predefining a grid
210 and allocating the flares to this grid. By using the source code written in R (R Core Team, 2013)



211 delivered by this study, the user can define the grid size independently. For calculating the average
212 over several days, the emissions for every single flare per day are calculated and summed up
213 according to their belonging to a certain grid box. This leads to one big point source per grid box. The
214 corresponding emissions are then averaged over the time period of interest for every grid box (flare
215 box hereafter). Considering this approach within an atmospheric model, by selecting the same grid
216 configuration for the flaring emission data and the model, no loss of information occurs.

217

218 3. Results

219

220 3.1 Spatial distribution of gas flaring in SWA

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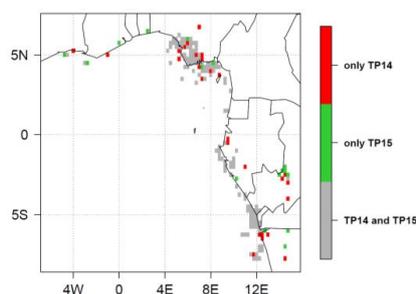
222 We have selected the two time periods June/July 2014 (TP14) and June/July 2015 (TP15) and omitted
223 all days without observations or with insufficient data coverage for VNP over SWA. This leads to 58
224 (48) observations for TP14 (TP15).

225 In the preparation of this work we have compared the estimated mean locations of the flares of TP14
226 with the Google Earth imagery (Google Earth, 2014) (not shown). Only the onshore flares are visible
227 in Google Earth. This visual verification reveals that 72% of the VNP detected onshore flares are
228 visible in Google Earth. It is very likely that the hit rate is much higher since it is often the case that
229 the Google Earth image quality is not good enough for verification or the images are not up to date.

230 This comparison indicates that VNP is a valid method to identify the flares in SWA.

231 For the following analysis we have calculated the emissions for both time periods on a grid with a
232 mesh size of 0.25° (28 km) from 10°S to 10°N and from 10°W to 15°E. Fig. 1 emphasizes the areas in
233 which VNP detects flares only in TP14 (TP15) in red (green) color and in grey the areas with flaring in
234 both periods.

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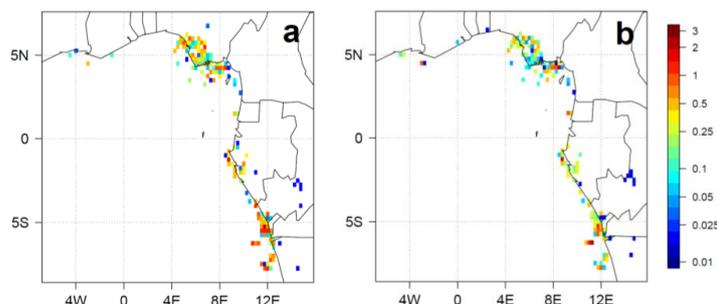
238 **Fig.1.** Flaring area for TP14 and TP15. Red (green) boxes denote areas with flaring only for TP14 (TP15). For the grey areas,
239 flaring is detected in both time periods.

240

241 Remarkable are the dominating flaring areas in the Niger Delta and the adjacent offshore regions in
242 the Gulf of Guinea. Also in the coastal region of Gabon, Republic of the Congo, Angola and
243 sporadically along the coast of Ivory Coast, Ghana and Benin, flaring occurs. By comparing TP14 and
244 TP15 more red than green areas are visible, especially in southern Nigeria, which indicates a
245 reduction in the flaring area from 2014 to 2015. A decrease in CO₂ from 1994 to 2010, particularly in
246 the onshore platforms is indicated by Doumbia et al. (2014).

247 The mean active flare density, as the sum over all detected flares in a box averaged over the time
248 period, is shown in Fig. 2 for (a) TP14 and (b) TP15.

249



250
251 **Fig.2.** Mean active flare density (number of active flares per box) averaged over (a) TP14 and (b) TP15 in logarithmic scale
252

253 Fig. 2 shows to a a reduction in the active flare density in TP15 compared to TP14. 72% of the flaring
254 area which TP14 and TP15 have in common, shows a reduction in TP15 about 48% on average. 28%
255 of the common flaring area shows an increase in TP15 about 124%. Therefore it seems that the
256 flaring intensity decreases in TP15 over large areas but simultaneously some flaring hotspots
257 occurred, which are distributed along the SWA coast (not shown). Fig. 2, together with the variation
258 of flaring emissions from TP14 to TP15 in Section 3.3.3, indicates the high year to year variations. This
259 makes the use of past averaged conditions questionable, especially when certain episodes are
260 studied.

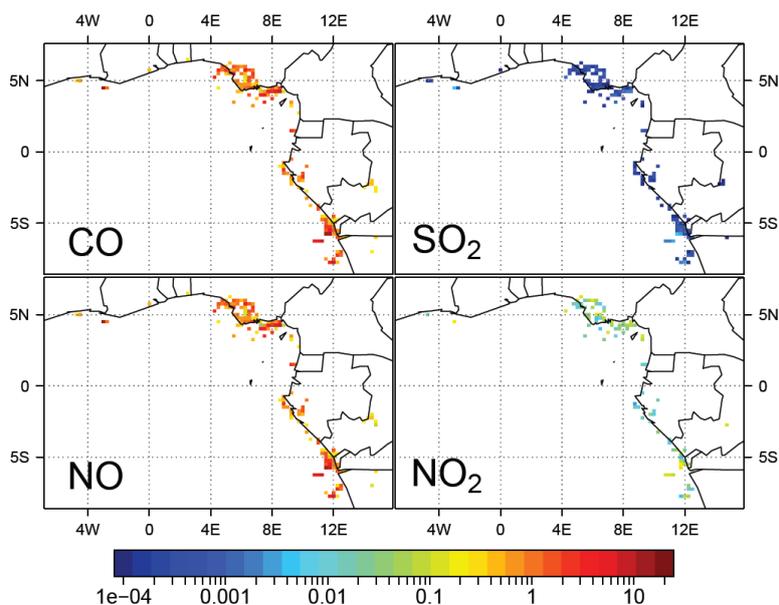
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262 3.2 Emission estimation

263

264 For the emission estimation we have used a climatological approach (E_{clim}). For every day with valid
265 data in TP14 and TP15 the emissions for all detected flares are calculated separately and allocated to
266 the predefined grid. The emissions are summed up in every flare box to have one joined flare per grid
267 box. Finally the temporal average for every grid box is calculated over TP14 and TP15 respectively.
268 Therefore all flares, detected in the time period, are active at once with their mean emission
269 strength. This method has the advantage that most likely all flares in the domain are captured even if
270 a fraction of them is covered by clouds at certain days. However, this could lead to an emission
271 overestimation because not all available flares are active at once. This problem of separating
272 between flares which are not active and flares which are active but covered by clouds and therefore
273 not visible for VNP is picked up again in Section 3.3.1. Fig. 3 shows the emissions of CO, SO₂, NO and
274 NO₂ in t h⁻¹ for TP15.

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276
277

Fig.3. Flaring emissions for TP15 within E_{clim} in t h^{-1} for CO , SO_2 , NO and NO_2

278
279

280 Highest emissions are calculated for carbon monoxide, followed by nitrogen oxide and nitrogen
281 dioxide. Sulfur dioxide shows lowest emissions since these emissions do not depend on combustion
282 processes but only on the natural gas composition (see Tab. 3) and the amount of flared gas (IU14).
283 Due to the use of the averaged measurements of Sonibare and Akeredolu (2004), local variations of
284 hydrogen sulfide concentrations in the natural gas cannot be taken into account. Hydrogen sulfide is
285 the only source of sulfur in the flared gas and therefore determines the emission of sulfur dioxide. To
286 assess this uncertainty, a sensitivity study with different hydrogen sulfide concentrations is given in
287 Section 3.3.5.

288

289 3.3 Estimation of uncertainties

290

291 In the following section the most relevant uncertainties are presented, together with approaches for
292 their assessment. This includes the uncertainty concerning the flare detection in the presence of
293 cloud cover, the uncertainty in the determination of the emitted heat flow H via the fraction of
294 radiated heat f , the uncertainty in the choice of the IU14 parameters and the changes in flare
295 operation from one year to another as well as the influence of the spatial variability of hydrogen
296 sulfide in the natural gas on the sulfur dioxide emissions. Apart from Section 3.3.4 all uncertainty
297 estimations are confined to TP15.

298

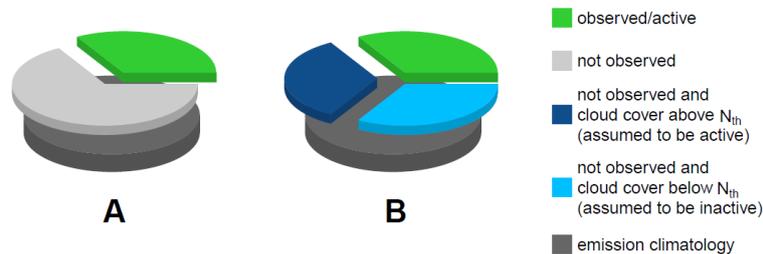
299 3.3.1 Uncertainty due to cloud cover

300

301 In Section 3.2 a climatological data set of flaring emissions (E_{clim}) was derived. When using this data
302 set we are losing the day to day variation of the flaring emissions that is delivered by VNP. Although
303 daily satellite observations are available, the problem arises that usually parts of the scene observed
304 by the satellite are covered by clouds. In the following we will describe a method of how to derive

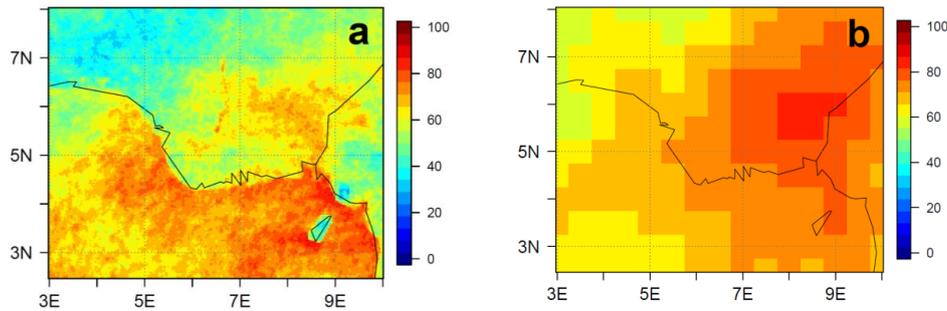


305 daily flaring emissions based on the climatological emissions (E_{clim}), a threshold of cloud coverage
 306 (N_{th}), and the actual detected flares at a certain day. This is illustrated schematically by Fig. 4.
 307 The closed grey pie in the lower layer of Fig. 4A gives the climatological number of flaring boxes in
 308 the research domain. At a certain day only within the green flaring boxes active flares are detected
 309 by VNP. The flaring boxes that are indicated in grey are those at which no active flares were detected
 310 by VNP, either because they are inactive or obscured by clouds. We now further separate this grey
 311 area by introducing an empirical threshold value N_{th} of cloud cover. In areas that belong to the grey
 312 fraction in Fig 4A, where the cloud cover is above N_{th} , we assume that the flares boxes are active and
 313 emit with their climatological emission values (since there are no current observations available).
 314 Those flare boxes are indicated by the dark blue color in Fig 4B. The light blue area indicates flare
 315 boxes where the cloud cover is below N_{th} and where no flares are detected by VNP. For this area we
 316 postulate that all flare boxes are inactive and consequently have zero emissions. Finally we calculate
 317 the total emissions at a certain day for $N_{\text{th}}=50\%$ (E_{50}), 75% (E_{75}) and 90% (E_{90}) as the sum of the
 318 climatological emissions in the dark blue area and the directly detected flares in the green area.
 319



320
 321
 322 **Fig.4.** Pie charts illustrating the flaring emission uncertainty assessment due to cloud cover for TP15. The entirety of the
 323 flare boxes within the emission climatology (E_{clim}) is given as closed grey pie in the bottom of **A** and **B**. **A** distinguishes
 324 between flare boxes in which flares are detected at a certain day (green) and the complement of undetected flare boxes
 325 (light grey). In **B** the light grey slice of **A** is separated in a cloud-covered (above cloud cover threshold N_{th} , dark blue) and
 326 cloud-free (below N_{th} , light blue) by using remote sensing observations. Flare boxes which are not detected by VNP and
 327 simultaneously show a cloud cover above N_{th} , are taken as active. Flare boxes which are not detected by VNP and
 328 simultaneously show a cloud cover below N_{th} , are taken as inactive. For N_{th} the values 50%, 75% and 90% are used. The
 329 higher N_{th} the smaller the dark blue slice in **B**.

330
 331 To separate the light grey slice in Fig. 4A in covered and uncovered flare boxes, we used
 332 instantaneous cloud fractional cover (CFC) from the geostationary Meteosat Second Generation 3
 333 (MSG3) (CM SAF, 2015, copyright (2015) EUMETSAT) for every day of TP15 around the time of VNP
 334 observation (Suomi-NPP overflight approx. at 1 UTC). This method is applied to all days of TP15 for
 335 every flare box.
 336 To ensure a consistent timing between cloud observation and VNP observation, the spatial domain
 337 was reduced with a focus on the Niger Delta area (see Fig. 5a) and the flares were allocated
 338 according to the cloud data grid with a mesh size of 0.03° .
 339

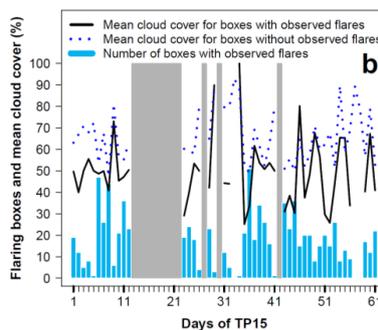


340
 341 **Fig.5.** Fractional cloud cover (%) observed from (a) the geostationary MSG3 and (b) the sun-synchronous Aqua/AIRS,
 342 averaged over TP15 around the time of VNP observation (approx. 1 UTC).
 343

344 Fig. 5a shows that the onshore flaring area for TP15 is in mean covered with clouds by 50-70%. For
 345 the offshore flaring area it is even higher with 70-90%. Therefore it is very likely that flares are
 346 frequently masked by clouds and therefore not detected by VNP. However, we suspect that the
 347 MSG3 cloud product underestimates (overestimates) the onshore (offshore) cloud cover when
 348 comparing with the findings of van der Linden et al. (2015). The high offshore coverage and the
 349 distinct land-water separation might be caused by overestimating low clouds in the presence of a
 350 warm and moist tropical ocean.

351 Fig. 5b shows a cloud climatology using Aqua/AIRS Nighttime data (Mirador, 2016). The Aqua/AIRS
 352 climatology shows higher cloud cover over land and no distinct separation between water and land
 353 surface. Both products identify the highest onshore cloud cover in the northeast of Port Harcourt
 354 (4.8°N, 7.0°E) and have similar values in the Nigerian offshore region (containing the offshore flares)
 355 of about 70-80%. The major difference in the climatologies appears onshore between 4.5°N and 6°N.
 356 This area includes the majority of the Nigerian onshore flares. Although it is not the aim of this study
 357 to identify the most reliable cloud climatology for SWA, it has to be considered that MSG3 likely
 358 underestimates the mean cloud cover over the Nigerian onshore flares up to 30%.

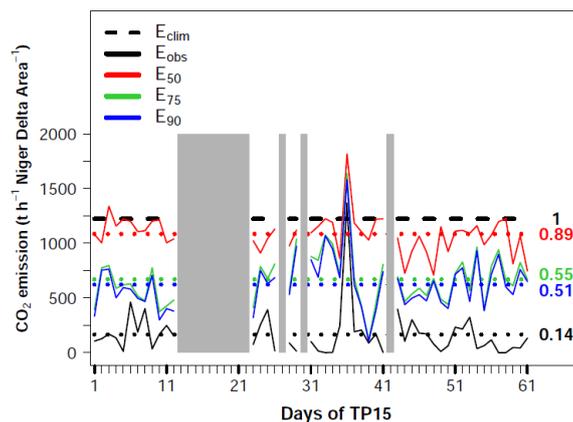
359 However, in the following the cloud climatology derived from MSG3 (Fig. 5a) is used since Aqua/AIRS
 360 cannot provide the full spatial coverage for every day (due to the sun-synchronous orbit of
 361 Aqua/AIRS).
 362



363
 364 **Fig.6.** Number of boxes with detected flares per day (blue bars) and the mean fractional cloud cover for the boxes with
 365 (without) detected flares as black solid (blue dotted) line (using MSG3, compare Fig. 5a). For the calculation of the latter,
 366 the cloud cover of the non-active flare boxes within E_{clim} are averaged (compare Niger Delta area in Fig. 2b). The grey
 367 shaded areas are omitted due to lack of VNP observation.
 368
 369



370 Fig. 6 shows the number of grid boxes with active flares per day in TP15 as blue bars. The grey areas
 371 indicate data gaps in VNP. E_{clim} includes 185 flare boxes according to the domain in Fig 5a. For TP15
 372 not more than 51 flare boxes are detected at once. In average only 8% of the total flaring area is
 373 active at once. As expected the temporal evolution of the flare boxes and the cloud cover for these
 374 boxes (black solid line in Fig. 6) shows an anticorrelation. The highest number of flare boxes at day 36
 375 is reached in a period of a comparatively low cloud cover. The mean cloud cover for the non-active
 376 flare boxes of E_{clim} (blue dotted line in Fig. 6), is in general higher than for the active flare boxes which
 377 implies that the cloud cover reduces the VNP detections. Fig. 6 also reveals that it is not suitable to
 378 use the strict cloud-free condition for the separation in Fig. 4B because nearly all of the boxes would
 379 be assigned to the dark blue cloud covered fraction and the resulting emissions would be nearly the
 380 same as E_{clim} .
 381 However, it has to be considered that the light points of flares are extremely small-scale signals
 382 (1/5000 of the VNP pixel, Zhang et al. (2015)) and even for an almost completely closed cloud deck
 383 VNP detections are possible.
 384 The climatology E_{clim} is the reference for this study. In addition we define E_{obs} which only considers
 385 the actually observed flares per day. E_{50} is defined as the combination of actually observed flares and
 386 cloud covered flares (see Fig. 4) with a cloud cover threshold of 50%. E_{75} (E_{90}) is equal to E_{50} but uses
 387 a cloud cover threshold of 75% (90%).
 388 To emphasize the difference between the different emission estimates, Fig. 7 shows the daily
 389 emissions of CO_2 for TP15 as a spatial sum over the Niger Delta area (see Fig. 4a). In contrast to E_{clim}
 390 (black dashed line), E_{obs} , E_{50} , E_{75} and E_{90} (solid lines) have a temporal variation within TP15.
 391



392
 393
 394 **Fig.7.** Daily CO_2 emissions (kg h^{-1}) within TP15 from flaring summed up over the Niger Delta area defined in Fig. 4a for the
 395 five emission estimates: E_{clim} (climatology, black dashed line), E_{obs} (VNP observations, black solid line), E_{50} (combination
 396 of VNP observations and the climatology for a cloud cover threshold of 50%, red solid line), E_{75} (as E_{50} but for a cloud cover
 397 threshold of 75%, green solid line), E_{90} (as E_{50} but for a cloud cover threshold of 90%, blue solid line). The dotted lines
 398 denote the spatiotemporal average of E_{obs} , E_{50} , E_{75} and E_{90} . The numbers on the right hand side show the ratios of the
 399 spatiotemporal averages E_{obs} , E_{50} , E_{75} and E_{90} towards E_{clim} . The grey shaded areas are omitted due to lack of VNP
 400 observation.

401
 402 E_{clim} delivers a daily CO_2 emission of about 1250 t h^{-1} within the Niger Delta area. The pure daily VNP
 403 observations within E_{obs} (black solid line) show only 14% of E_{clim} emissions (numbers on left hand side
 404 of Fig. 7) on average (black dotted line). The emissions from VNP observations together with the
 405 climatology for the cloud threshold of 50% within E_{50} (red solid line) is closest to the climatology (89%



406 of E_{clim} , red dotted line). The high overall cloud cover within the domain (compare with blue dotted
 407 line in Fig. 6) together with the relative low cloud cover threshold leads to the result, that nearly the
 408 complete climatology is used for E_{50} and therefore the difference to E_{clim} is small. The emissions from
 409 VNP observations together with the climatology for the cloud threshold of 75% and 90% within E_{75}
 410 and E_{90} (green and blue solid line) shows only small deviations but are significantly reduced in
 411 comparison to E_{clim} (55% and 51% of E_{clim} , green and blue dotted line). Day 36 of TP 15 shows highest
 412 emissions in E_{obs} , E_{50} , E_{75} and E_{90} , owing to the combination of low cloud cover and high flaring
 413 activity (compare with Fig. 6). Regarding the uncertainty in the cloud cover climatology (compare Fig.
 414 5a and Fig. 5b), the emissions of E_{50} , E_{75} and E_{90} might be underestimated. The underestimation of
 415 the cloud cover in the onshore flaring area could lead to an unjustified increase in flare boxes below
 416 N_{th} and therefore to a reduced number of active flares per day.

417 These emission estimations contain different information. E_{clim} includes all flares of the domain
 418 despite cloud cover but can overestimate the emissions. E_{obs} shows the VNP reality, including a
 419 temporal development, but cannot consider the cloud-covered flares. E_{50} , E_{75} and E_{90} combine the
 420 flare location information of E_{clim} and the full temporal resolution of VNP in E_{obs} by using cloud
 421 observations. However this approach is based on the assumption that all cloud covered flare boxes
 422 are active, which is also linked to high uncertainty. Additionally E_{50} , E_{75} and E_{90} depend on the
 423 availability of a longer VNP observational dataset. The ratios of the spatiotemporal means of E_{obs} , E_{50} ,
 424 E_{75} and E_{90} to the spatial mean of E_{clim} (as denoted by the numbers in Fig. 7) are used as correction
 425 factors (CF) for E_{clim} in the following (see Tab. 4). E_{clim} is taken as the reference (CF=1).

426

427 **Tab.4.** Emission estimations including information about flaring (daily observation and climatology) and cloud cover
 428 observation. The correction factors (CF) are derived for TP15 from a spatiotemporal emission mean in the Niger Delta area
 429 (2.5°N-8°N, 3°E-10°E) and refer to E_{clim} .

430

Name	Emission estimate	CF for E_{clim}
E_{clim}	Climatology (reference)	1
E_{obs}	Observed flares	0.14
E_{50}	Observed flares + climatology ($N_{\text{th}} = 50\%$)	0.89
E_{75}	Observed flares + climatology ($N_{\text{th}} = 75\%$)	0.55
E_{90}	Observed flares + climatology ($N_{\text{th}} = 90\%$)	0.51

431

432 These CF are a simple method to include the information of E_{obs} , E_{50} , E_{75} and E_{90} into E_{clim} by
 433 multiplying E_{clim} with the corresponding correction factor. In this case the same 185 flare boxes of
 434 E_{clim} are used but with an emission strength reduced to the averaged conditions of E_{obs} , E_{50} , E_{75} and
 435 E_{90} . This approach is based on the assumption that the correction factor, deduced for the Niger Delta
 436 area, is valid for the whole domain specified in Section 3.1. This assumption seems to be justified
 437 since the Niger Delta area contains most of the gas flares in the domain.

438

439 3.3.2 Uncertainty due to IU14 input parameters

440

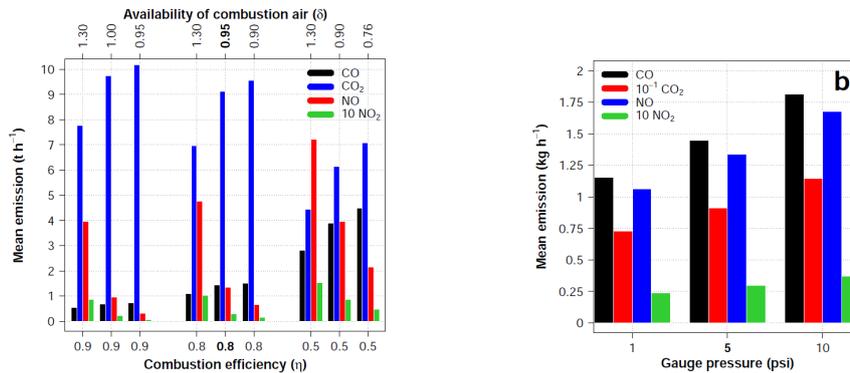
441 To assess the uncertainty which arises from the combustion efficiency η and the availability of
 442 combustion air δ , a sensitivity study has been carried out. The exact values for the SWA flares are
 443 unknown and very likely highly variable from one flare to another, depending on the flare type and
 444 operation. Fig. 8a shows the flare emissions averaged over SWA and TP15 for CO, CO₂, NO and NO₂.
 445 The parameters η and δ are varied referring to IU14. A complete combustion ($\eta = 1$) does not
 446 produce CO emissions since all carbon is transformed to CO₂ (not shown). With decreasing η and δ ,
 447 the CO and CO₂ emissions increase. Concerning CO we assume the lower limit for $\eta = 0.9$ and



448 $\delta = 1.3$ (left of Fig. 8a) and the upper limit for $\eta = 0.5$ and $\delta = 0.76$ (right of Fig. 8a). The values
 449 used for this study are located in the center of Fig. 8a. By taking the latter as reference, the lower
 450 (upper) limit leads to a decrease (increase) in CO emission of -63% (+210%). For CO₂ we derived an
 451 upper (lower) limit of +38% (-72%).

452 A higher combustion efficiency or a higher availability of combustion air allows an enhanced
 453 formation of NO. Therefore NO emissions increase (decrease) with decreasing η (δ). We assume the
 454 lower limit for $\eta = 0.9$ and $\delta = 0.95$ and the upper limit for $\eta = 0.5$ and $\delta = 1.30$. Taking again the
 455 central parameter set of Fig. 8a as reference, the lower (upper) limit leads to a decrease (increase) in
 456 NO emission of -77% (+441%).

457



458 **Fig.8.** Flaring emissions (kg h^{-1}) spatiotemporally averaged over SWA and TP15 depending on (a) combustion efficiency η
 459 and availability of combustion air δ and (b) gauge pressure (psi) for the setup of η and δ which is used for this study
 460 (emphasized in bold). SO₂ is not shown because it does not depend on η or δ .
 461

462 The emissions of NO₂ are comparatively low owing to the source temperature which is in general
 463 lower than the NO₂ formation threshold of 1600 K.

464 In addition, Fig. 8b shows the emissions depending on the gauge pressure for 1 (lower limit), 5 and
 465 10 psi (upper limit) (7, 34 and 69 kPa respectively) for $\eta = 0.8$ and $\delta = 0.95$. Regarding 5 psi as the
 466 reference, the lower (upper) limit leads to a decrease (increase) in CO emissions of -21% (+26%).

467 Fig. 8 emphasizes that the technical conditions of flaring crucially influence the emission strength and
 468 that the emissions are more sensitive towards η and δ than towards the gauge pressure.
 469

470

471 3.3.3 Uncertainty due to the fraction of radiated heat

472 To estimate the uncertainty in the fraction of radiated heat f (see Tab. 2), we have used the standard
 473 deviation of the literature values given in the appendix of Guigard et al. (2000) in addition to the
 474 mean value of $f = 0.27$. This leads to a domain of uncertainty for the value f of $(^{0.38}/_{0.16})$. Therefore
 475 the VNP observed radiant heat is multiplied with the factor $1/f$ of $3.7 (^{6.2}/_{2.6})$.
 476

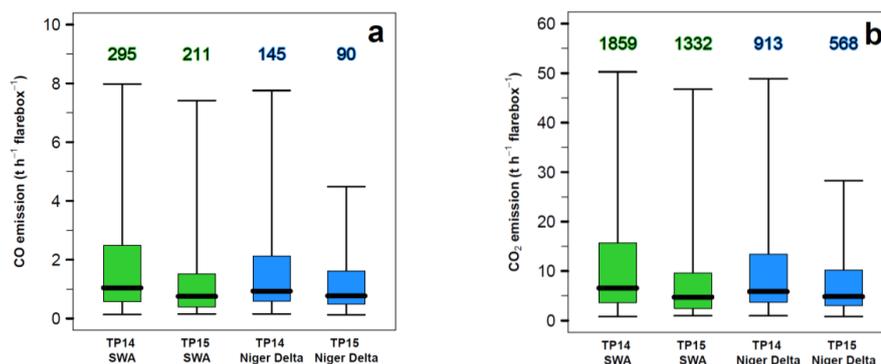
477

478 3.3.4 Interannual variability

479 The differences in flaring between TP14 and TP15, indicated in Fig. 1 and Fig. 2, are quantified in this
 480 section according to the emissions of CO (Fig. 9a) and CO₂ (Fig. 9b). The boxplots include all flaring
 481 boxes for the two domains SWA (green) and the Niger Delta area (blue). The numbers above indicate
 482 the integrated emissions per hour and area in tons.
 483



484



485

486 **Fig.9.** Flare box emissions of (a) CO and (b) CO₂ (E_{clim} , t h⁻¹ flarebox⁻¹) for SWA (green) and the Niger Delta area (blue) for
487 TP14 and TP15. The values above the boxplots indicate the emissions per hour, integrated over SWA (green) and the Niger
488 Delta area (blue). The whiskers span the data range from the 0.025-quantile to the 0.975-quantile (95% of the data). Data
489 outside of this range is not shown.

490

491 The emissions of CO₂ are 6.3 times higher than the CO emissions. For SWA the mean value of
492 emissions is statistically significant lower for TP15 compared to TP14 (Wilcoxon-Mann-Whitney rank
493 sum test with a significance level of 0.05). For the Niger Delta area the emission means show no
494 significant difference. The significant different mean values for SWA emissions emphasize the
495 relevance of using a flaring dataset which is up to date to reduce uncertainties arising from
496 deviations in flare locations or flaring processes.

497

498 3.3.5 Uncertainty due to spatial variability in H₂S

499

500 Since hydrogen sulfide (H₂S) is the only sulfur source in the flared gas, it determines the emission of
501 sulfur dioxide. The natural gas composition measurements from the ten flow stations given in
502 Sonibare and Akeredolu (2004) contain only two stations with nonzero H₂S content. Therefore
503 averaging over the ten stations (see Tab. 3) leads to a low H₂S content in the emission calculations.
504 By using the highest concentration value of H₂S given in Sonibare and Akeredolu (2004) (see Tab. 3,
505 H₂S concentration 0.03% instead of 0.005%), we try to estimate the upper limit of SO₂ emission,
506 assuming that all flares are provided with this more sulfur containing gas. With this approach the
507 spatiotemporal averaged SO₂ emissions increase from 0.6 to 4.9 kg h⁻¹. The maximum values in the
508 flare boxes increase from 4.7 to 41.8 kg h⁻¹. These are rather low values.

509 This comparison reveals that among the flaring conditions also the natural gas composition plays an
510 important role in estimating the flaring emissions reasonably. To rely on a single measurement
511 dataset for a large flaring domain and without taking into account spatial variability is therefore
512 problematic but has to be accepted owing to insufficient data.

513 This section has estimated the uncertainties in gas flaring due to cloud cover, parameters of IU14, the
514 fraction of radiated heat, the temporal variability and the H₂S concentration in the natural gas. The
515 uncertainty regarding the spatial variability of the total hydrocarbon fraction of the natural gas,
516 which is estimated by the variations in the ten flow station measurements of Sonibare and Akeredolu
517 (2004), is below 1%.

518 However, there are further assumptions or sources of uncertainty which cannot be quantified within
519 this study: We assume that the natural gas composition, which is measured in one region, is valid for



520 SWA entirely. The gas flares are taken as constant emission sources because VNP only provides one
521 observation (overflight) per day. We cannot take into account the spatial variability of the flares
522 concerning the IU14 parameters and the stack heights. And finally IU14 delivers no VOCs and black
523 carbon.

524

525 4. Comparison with existing emission inventories

526

527 The following section places the estimated flaring emissions of this study in the context of existing
528 emission inventories, by taking the focus on CO₂. A direct comparison with existing emission
529 inventories is problematic due to different reference time periods, spatial domains, definitions of
530 emission sectors and the limitation of chemical compounds. Tab. 5 summarizes the CO₂ emissions for
531 different inventories regarding Nigeria or the Niger Delta area as denoted in Fig. 5a, the flaring
532 hotspot of the research domain. The results of this study shows no flaring in the northern part of
533 Nigeria and therefore flaring within the Niger Delta area can be seen as the total flaring area of the
534 country. To derive annual emission values for the results of this study, it is assumed that the flaring
535 emission conditions of TP14 and TP15 are representative for the whole year 2014 and 2015
536 respectively. Therefore the hourly emissions are integrated over 365 days.

537

538 **Tab.5.** Comparison between existing emission inventories for CO₂ (with a focus on gas flaring if available) and the results of
539 this study for Nigeria or the Niger Delta area in teragram (Tg) per year. For TP14 and TP15 it is assumed that the two month
540 observations represent the flaring conditions of the whole year 2014 and 2015 respectively. Therefore the emissions were
541 integrated to yearly values. The values in brackets represent the upper and lower limit owing to the uncertainties estimated
542 in Section 3. For the fraction of radiated f the mean value 0.27 and the lower (upper) boundary of 0.16 (0.38) are used,
543 representing a further source of uncertainty. The products given in bold are directly related to flaring emissions.

544

Emission inventory	Time period	CO ₂ emissions (Tg y ⁻¹)		
		$f = 0.16$	$f = 0.27$	$f = 0.38$
This study (E_{clim})	2014 (from TP14)	13 (¹⁹ / ₃)	8 (¹² / ₂)	6 (⁸ / ₁)
This study (E_{clim})	2015 (from TP15)	8 (¹² / ₂)	5 (⁷ / ₁)	4 (⁵ / ₀)
This study (E₇₅)	2014 (from TP14)	7 (¹¹ / ₂)	4 (⁷ / ₁)	3 (⁵ / ₀)
This study (E₇₅)	2015 (from TP15)	5 (⁷ / ₁)	3 (⁴ / ₀)	2 (³ / ₀)
CDIAC (2015b)¹	2011	27.47		
EIA (2015)²	2010; 2011; 2013	38.81; 41.39; 52.83		
Doumbia et al. (2014)¹	2010	45		
EDGAR 4.2³ (ECCAD, 2015)	2008	8.75		
EDGAR 4.2⁴ (ECCAD, 2015)	2008	3.50		
EDGAR 4.3.2 ⁵ (EDGAR, 2016)	2010; 2011; 2012	29.4, 28.8, 28.9		
EDGARv43FT2012 ⁶ (EDGAR, 2014)	2014	93.87		

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555 The CO₂ emission estimations of this study are given in Tab.5 together with an overall uncertainty
556 range (⁺³⁸/₋₇₂ %) including the uncertainty from the IU14 parameters η and δ (⁺¹²/₋₅₂ %) and the
557 gauge pressure (⁺²⁶/₋₂₁ %) and from spatial variability of total hydrocarbon. The latter uncertainty is
558 small (below 1%) owing to the low variation in THC concentration in the measurements of Sonibare



559 and Akeredolu (2004). The uncertainty due to cloud cover is represented by E_{75} . Regarding the
560 relatively large uncertainty there is no preference in one of the emission estimates E_{clim} and E_{obs} .
561 By assuming the uncertainty range of the fraction of radiated heat f between 0.16 and 0.38, the
562 results of the study on hand show CO_2 emissions in the same order of magnitude as the Carbon
563 Dioxide Information Analysis Center (CDIAC, 2015b), the Energy Information Administration (EIA,
564 2015) and the EDGARv.4.3.2 (EDGAR, 2016) database, with best results for $f = 0.16$ but with an
565 overall tendency to underestimate the emissions. E_{clim} shows smaller deviations to the existing
566 inventories than the cloud correction approach of E_{75} . A direct comparison is hindered by a time lag
567 of 3-4 years and missing information about the uncertainties of CDIAC. The values of EIA are higher
568 than those of CDIAC because EIA includes the consumption of natural gas in addition to gas flaring.
569 Doumbia et al. (2014) combines Defense Meteorological Satellite Program (DMSP) observations of
570 flaring with the emission factor method to derive flaring emissions. The results agree with EIA (2015)
571 but are 64% higher than CDIAC (2015b).

572 The emission inventory EDGAR v4.2 (ECCAD, 2015) delivers $8.75 (3.50) \text{ Tg CO}_2 \text{ y}^{-1}$ for Nigeria (Niger
573 Delta area) for the emission sector *refineries and transformation*, which is in good agreement with
574 the results for the study on hand.

575 As a benchmark for the flaring CO_2 , the total CO_2 emissions for Nigeria are given by EDGAR (2014),
576 (fossil fuel use and industrial processes). Taking EDGAR (2014) as a reference for total CO_2 emissions
577 of Nigeria, flaring emissions contributes by $9 (13/2)\%$ (2014; $E_{\text{clim}} f = 0.27$), $14 (20/3)\%$ (2014; E_{clim}
578 $f = 0.16$), 9% (2008; ECCAD, 2015), 28% (2011; CDIAC, 2015b), 48% (2010; Doumbia et al., 2014) or
579 56% (2013; EIA, 2015). The large spread between the different inventories emphasizes the large
580 uncertainty within the estimation of emissions from gas flaring.

581 A shortcoming of the PEGASOS_PBL-v2 (not shown) and the EDGAR v4.2 emission inventory is the
582 lack of offshore flaring emissions in the Gulf of Guinea south of Nigeria. For CDIAC and EIA this
583 cannot be verified since the data is only available as a single value per country.

584 The differences between the results of this study and the existing emission inventories might be
585 caused by an underestimation of the flow rate by VNP and Eq. 1 or by an inconsistent definition of
586 emission source sectors for the existing inventories. E_{clim} , E_{75} , Doumbia et al. (2014) and CDIAC
587 (2015b) focus on gas flaring, whereas other products also include natural gas consumption and
588 emissions from refineries and transformation which also can include non-flaring emissions within and
589 outside the areas indicated as flaring area by the satellite imagery. In addition, the existing
590 inventories do not provide current values (time lag of 2 to 6 years) and therefore not consider the
591 emission reduction indicated by Fig. 9.

592

593 **5. Discussion and conclusions**

594

595 The gas flaring emission estimating method of Ismail and Umukoro (2014) (IU14) has been combined
596 with the remote sensing flare location determination of the VIIRS Nightfire Prerun V2.1 Flares only
597 (VNP) (VIIRS, 2015) for a new flaring emission parameterization. The parameterization combines
598 equations of incomplete combustion with the gas flow rate derived from remote sensing parameters
599 instead of using emission factors and delivers emissions of the chemical compounds CO, CO_2 , NO and
600 NO_2 .

601 Within this study the parameterization was applied to southern West Africa (SWA) including Nigeria
602 as the second biggest flaring country. Two two-month flaring observation datasets for June-July 2014
603 and 2015 were used to create a flaring climatology for both time periods. In this climatology all
604 detected flares emit with their mean activity.



605 The uncertainties owing to missed flare observations by cloud cover, parameterization parameters,
606 interannual variability and the natural gas compositions were assessed. It can be shown that the
607 highest uncertainties arise from the definition of the fraction of radiated heat f and the IU14
608 parameters. By using remote sensing cloud cover observations, a correction factor for the flaring
609 climatological emission was derived which reduces the mean emissions about 50%. However, owing
610 to the large uncertainty ranges, no significant difference between the climatological inventory and
611 the cloud corrected inventory can be stated. Comparing the emissions of 2014 and 2015, a reduction
612 in the flaring area, density of active flares and a significant reduction in SWA emissions about 30%
613 can be observed, which underlines the need for more recent emission inventories.

614 The uncertainty due to the natural gas composition is compound dependent. The spatial variation in
615 total hydrocarbon is negligible but the availability of hydrogen sulfide, which exclusively determines
616 the amount of emitted SO_2 , cause large uncertainty. By taking the combustion efficiency to derive the
617 fraction of unburned natural gas, the amount of emitted VOCs might be estimated in addition to the
618 species of the study on hand but would also be linked to high uncertainties concerning the VOC
619 speciation. The uncertainty in VOC emission is increased drastically by natural gas which is vented
620 directly into the atmosphere instead of being flared, since the venting cannot be detected by VNP.

621 With a focus on Nigeria, the CO_2 emission estimates of this study were compared with existing
622 inventories. For the climatology, CO_2 emissions of $8 \text{ (}^{12}/_2\text{) Tg y}^{-1}$ for 2014 and $5 \text{ (}^7/_1\text{) Tg y}^{-1}$ for 2015
623 were derived. EDGAR v4.2 for the year 2008 shows the same order of magnitude when limiting to
624 emissions from refineries and transformation. CDIAC (Carbon Dioxide Information Analysis Center) is
625 in the same order of magnitude as the results of this study. Doumbia et al. (2014) and EIA (Energy
626 Information Administration) show emissions which are 2.4 and 2.8 times higher than the results of
627 this study. The deviations might be caused by uncertainties in the flow rate derived by VNP radiant
628 heat, which can be assessed only rudimentary via the parameter of the fraction of radiated heat.
629 Additionally, the usage of emission factors in the existing inventories which did not take into account
630 the spatiotemporal variability of flaring, inconsistent emission sector definitions or the time lag of
631 the emission inventories of 2-5 years can lead to deviations. The positive trend in Nigerian gas flaring
632 CO_2 emissions derived by EIA from 38.81 to 52.83 Tg y^{-1} between 2010 and 2013 contradicts the
633 findings of Doumbia et al. (2014) and this study, which generally show a decrease in emissions from
634 1994 to 2010 and from 2014 to 2015, respectively. Based on the sensitivity study, which reveals high
635 uncertainties of the flaring emission, we conclude that there is no preference in the choice of one of
636 the emission estimates presented in this study. Therefore we recommend the use of the
637 climatological approach when using the R package.

638 Despite the generally large uncertainties in the estimation of emissions from gas flaring, this method
639 allows a flexible creation of flaring emission datasets for various applications (e.g. as emission
640 inventory for atmospheric models). It combines observations with physical based background
641 concerning the combustion. The use of current data makes it possible to consider present trends in
642 gas flaring. Even the creation of near real-time datasets with a time lag of one day is possible. The
643 emissions are merged on grid predefined by the user and depending on the availability of VNP data,
644 the temporal resolution can be selected from single days to years.

645 An improvement of this parameterization can be achieved by an extension of the IU14 method to
646 black carbon and VOCs and an inclusion of spatial resolved measurements of the natural gas
647 composition in combination with information of the gas flaring processes from the oil producing
648 industry.

649



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651
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657 into our parameterization. We also thank the Earth Observation Group (EOG) of NOAA for providing
658 the *VIIRS Nightfire Flares Only* product.

659

660 **Code and/or data availability**

661 This publication includes a package of well documented R scripts which is free available for research
662 purposes and enables the reader to create their own gas flaring emission datasets. It includes
663 exemplarily the preprocessing for June-July 2015 with a focus on southern West Africa. You get
664 access to the code via zenodo.org (DOI: 10.5281/zenodo.50938).

665

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